

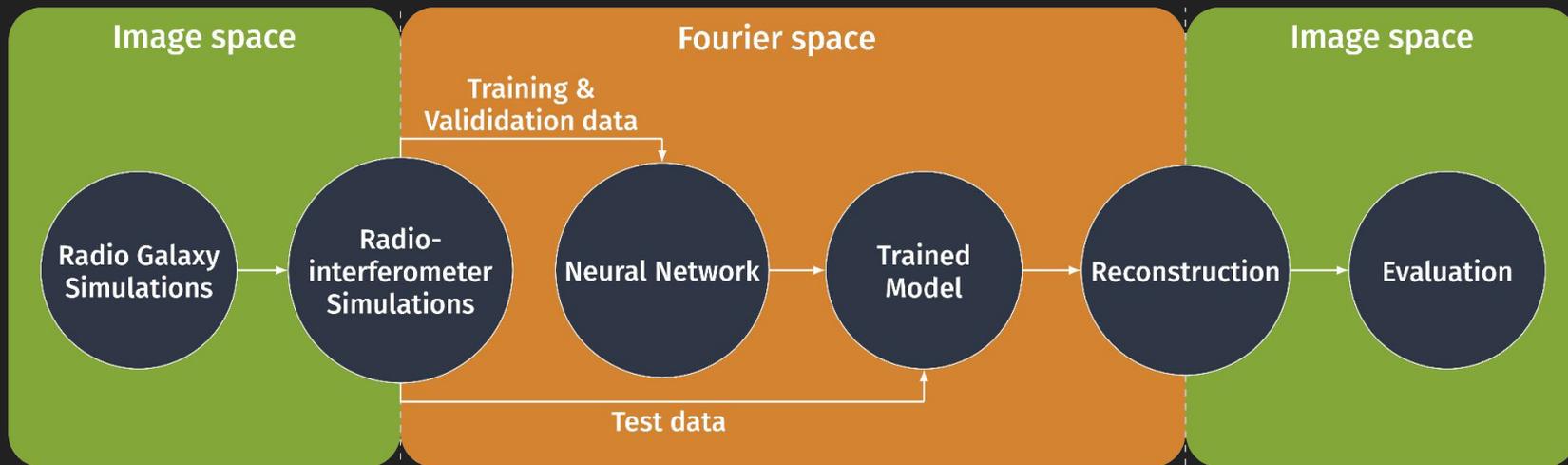
Deep learning-based imaging of Multi-source observations

Kevin Schmitz

TU Dortmund University
LAMARR Institute for ML and AI

15.12.2025

radionets-project



Simulations

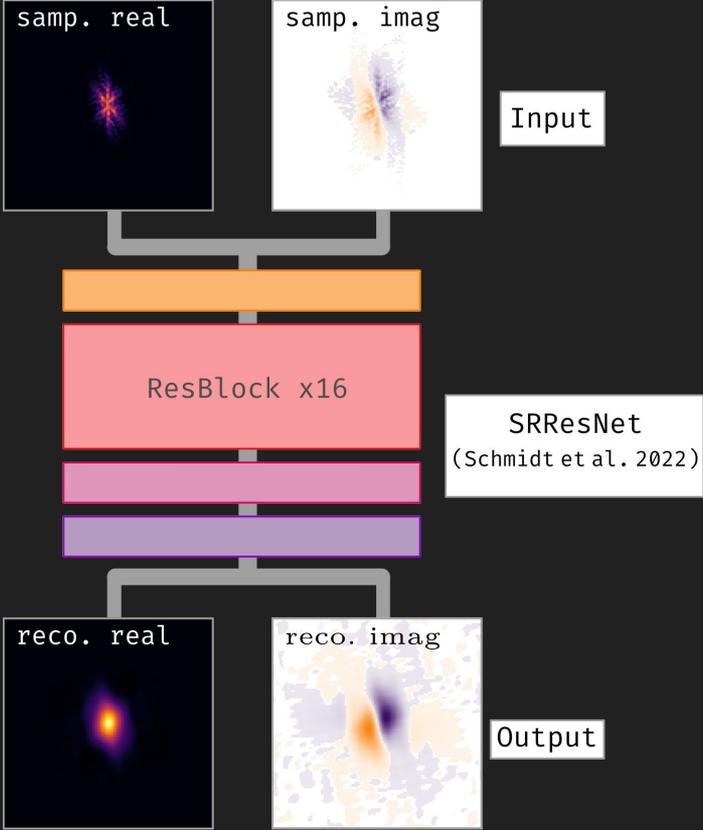


Model Building

Evaluation



Deep Learning-based Reconstruction



MeerKat data set

L- Band observations of Virgo cluster

pointing at the outer part

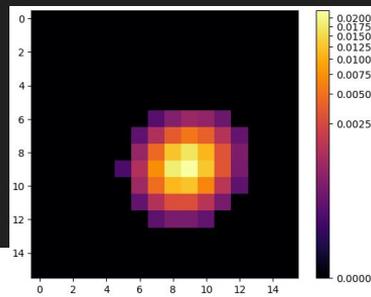
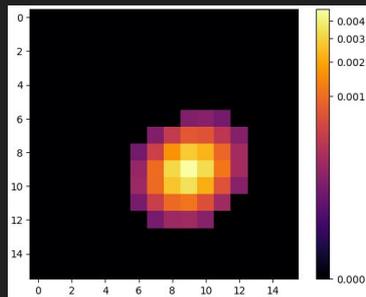
2 arcsec resolution

350.000 visibilities

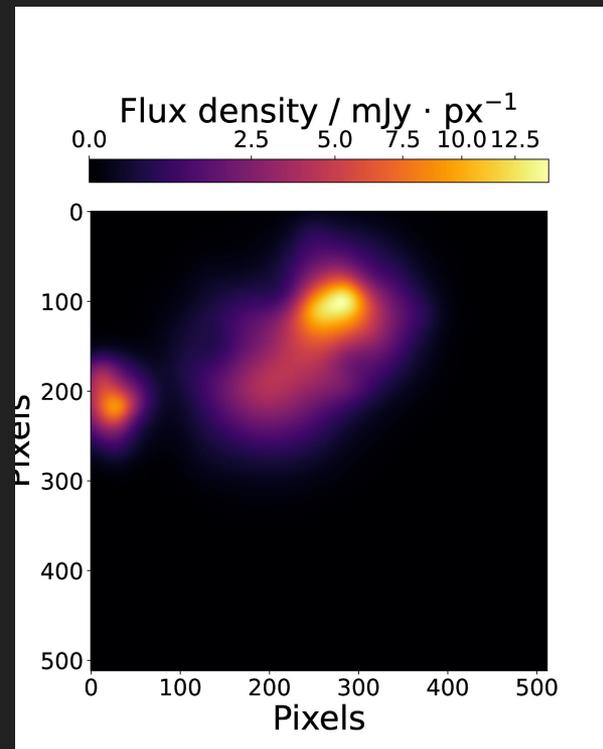
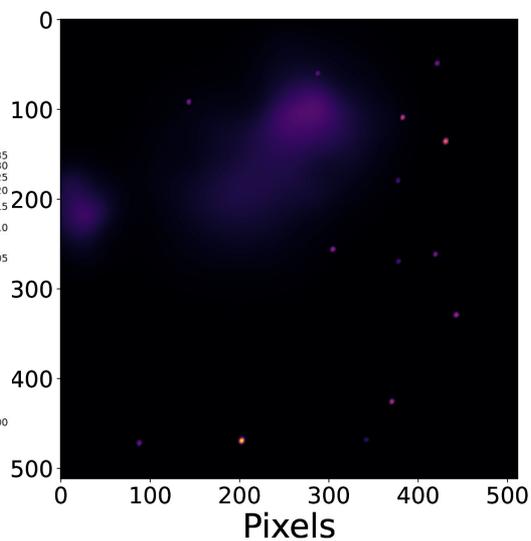
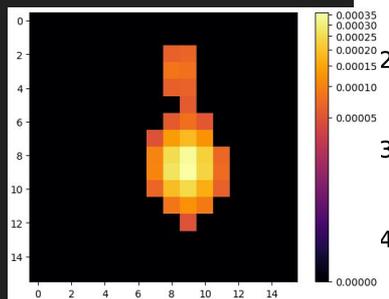
1.3 GHz – 16 channels



Synthetic radio skies



Flux density / mJy · px⁻¹

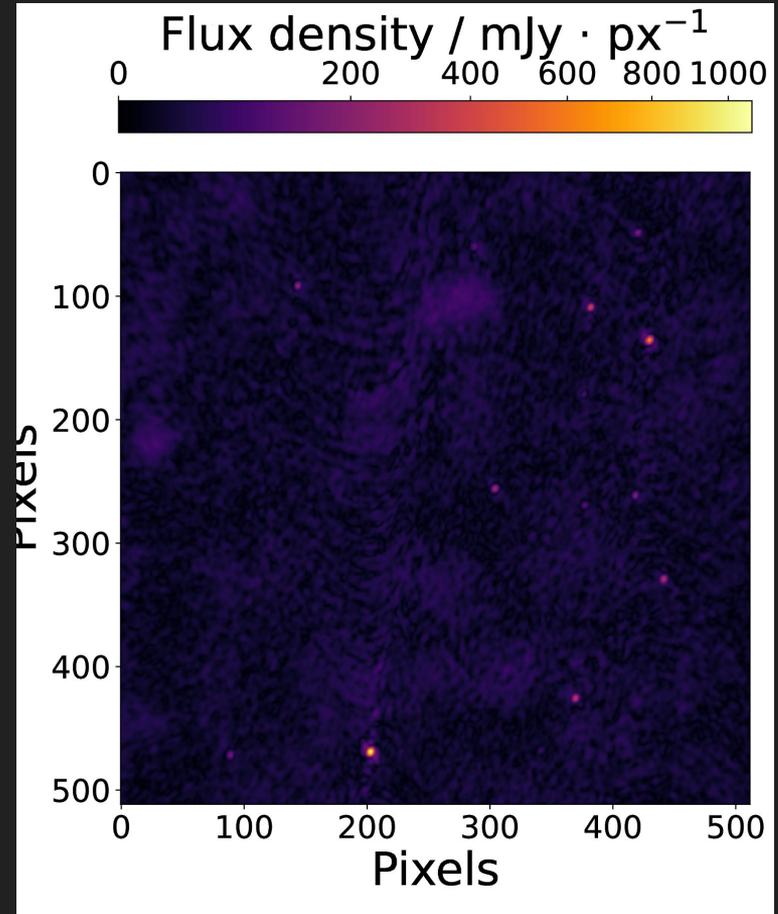


Visibility simulations

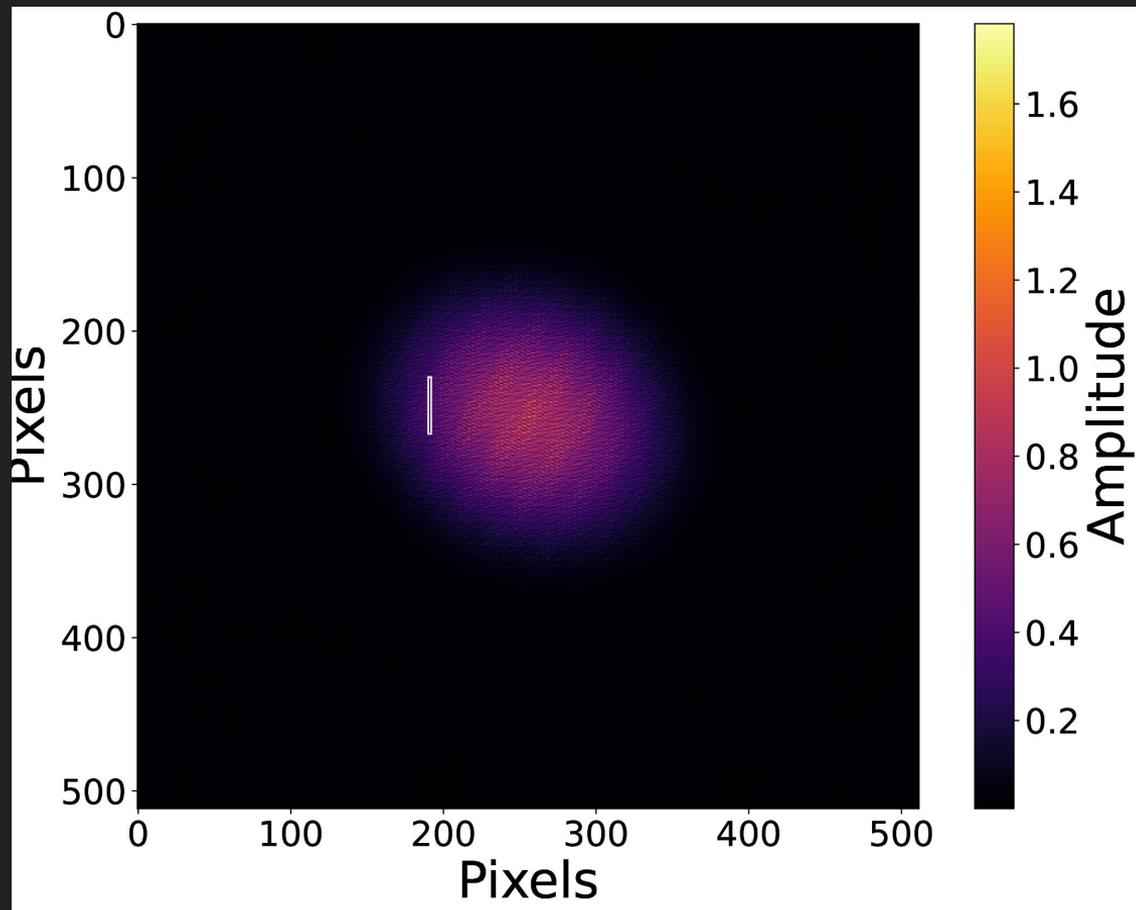
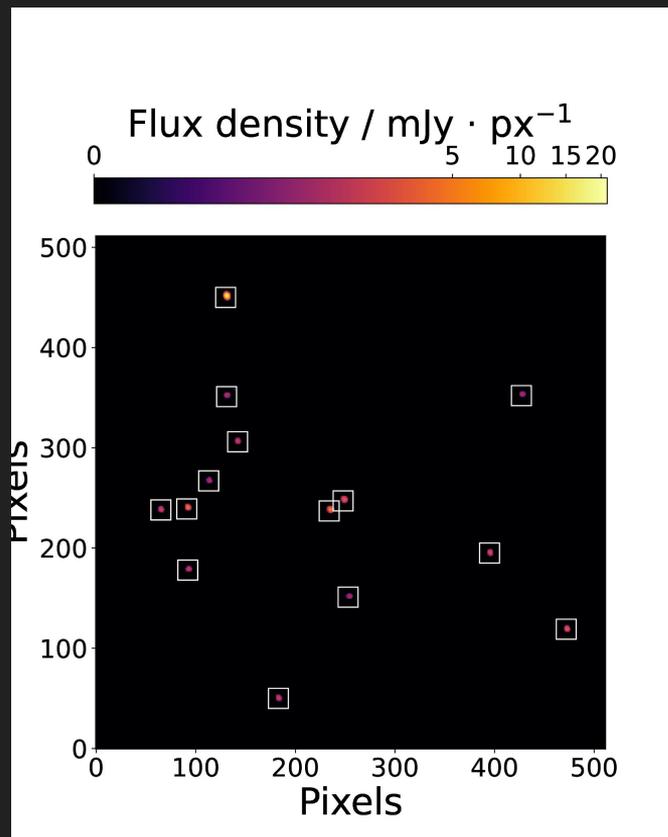
Radio interferometer measurement equation

Considering (u, v, w) and (l, m, n)
-> non-coplanar effects

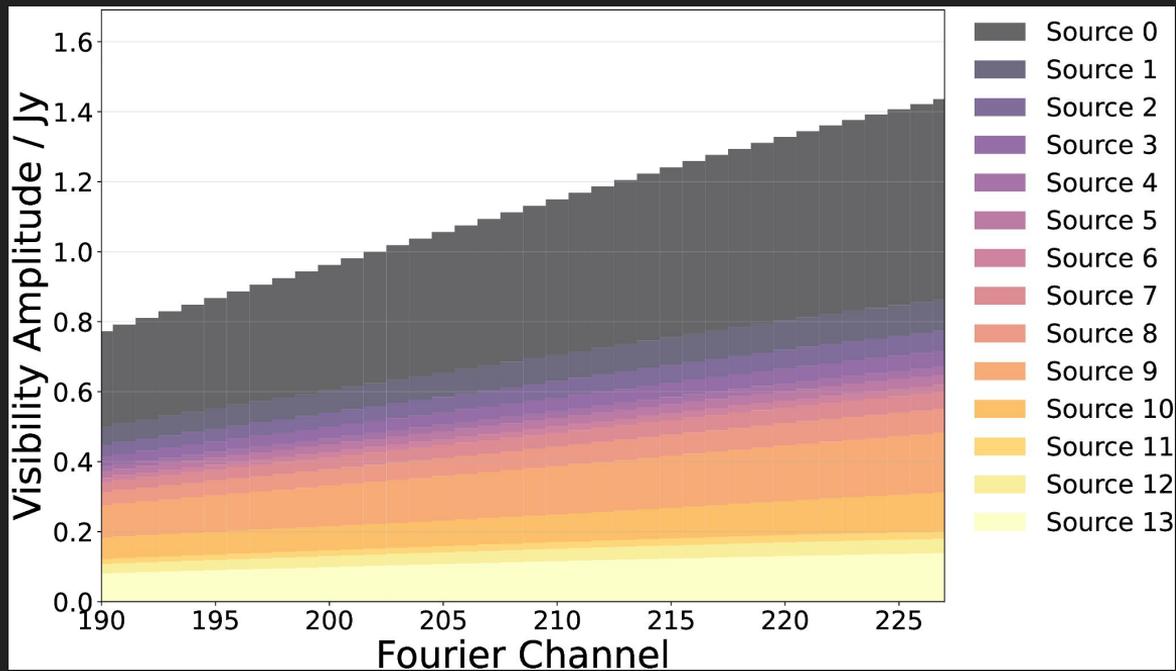
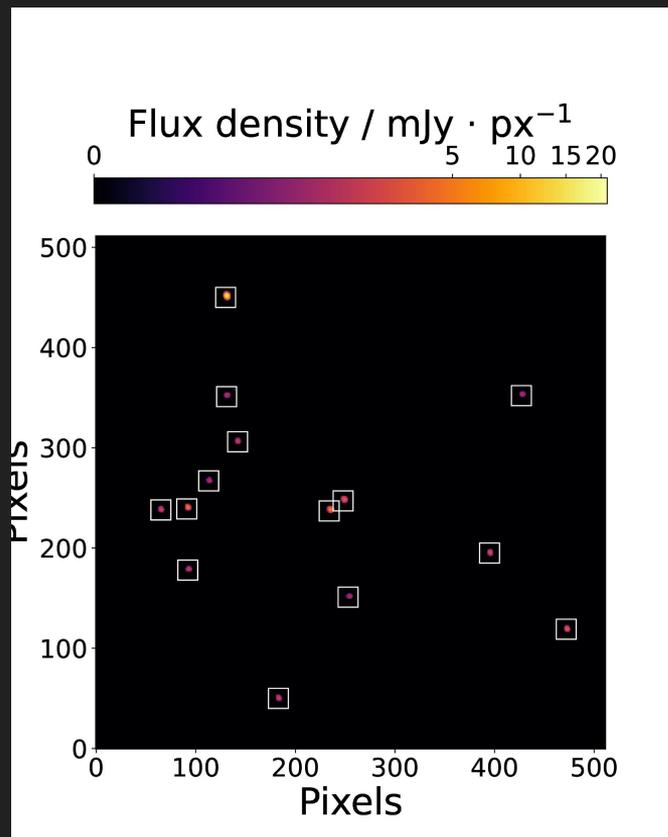
Import when creating dirty images



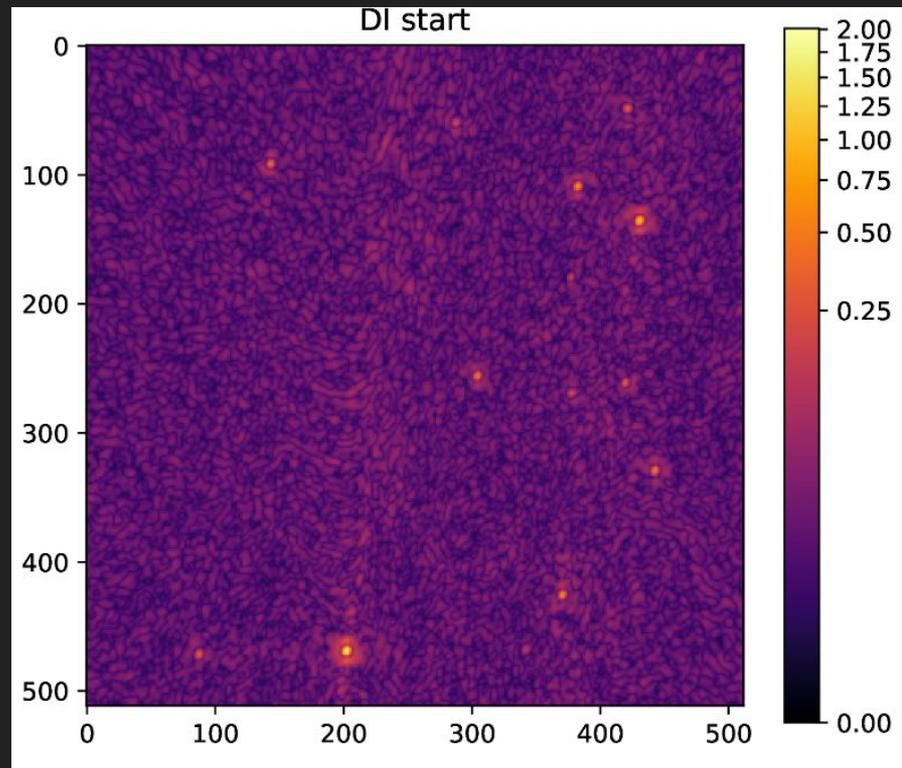
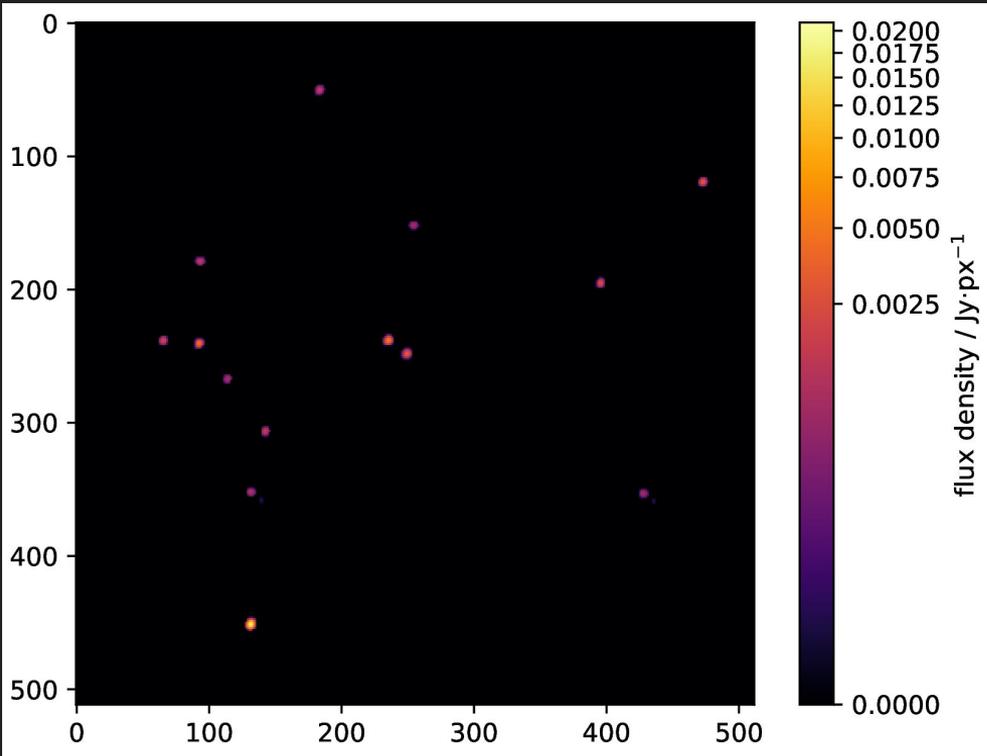
Visibility data



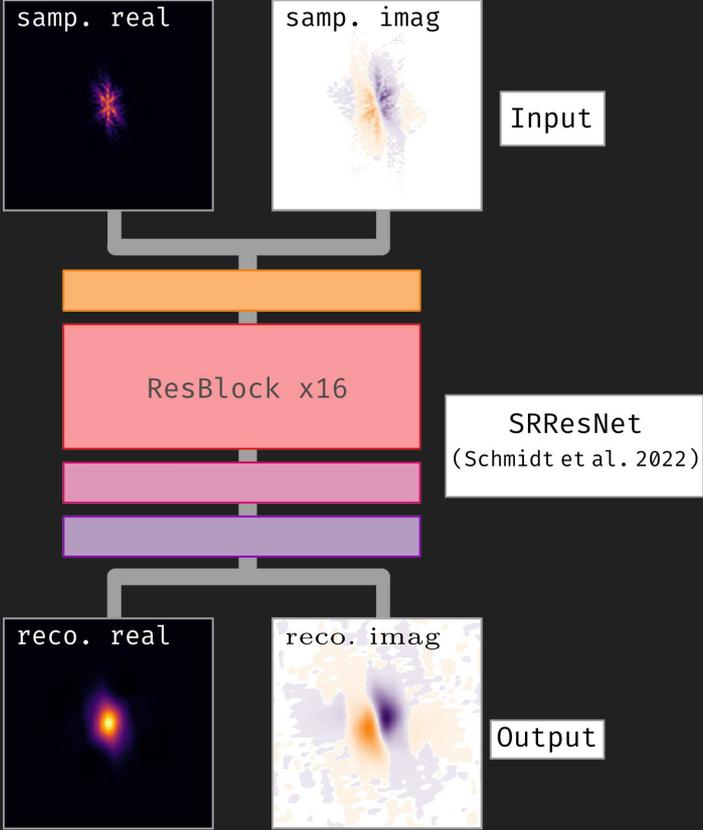
Visibility data



Visibility data



Deep Learning-based Reconstruction

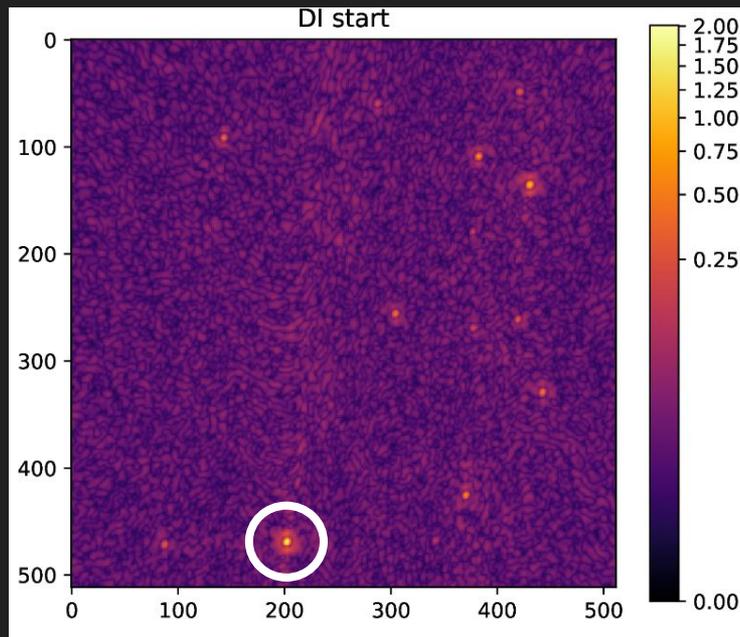


Process sky chunks

Create sky chunk with brightest source

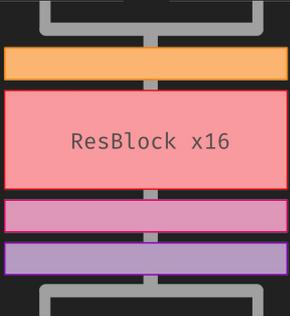
Reconstruction in visibility space with neural network

Subtract solution from complete visibility set

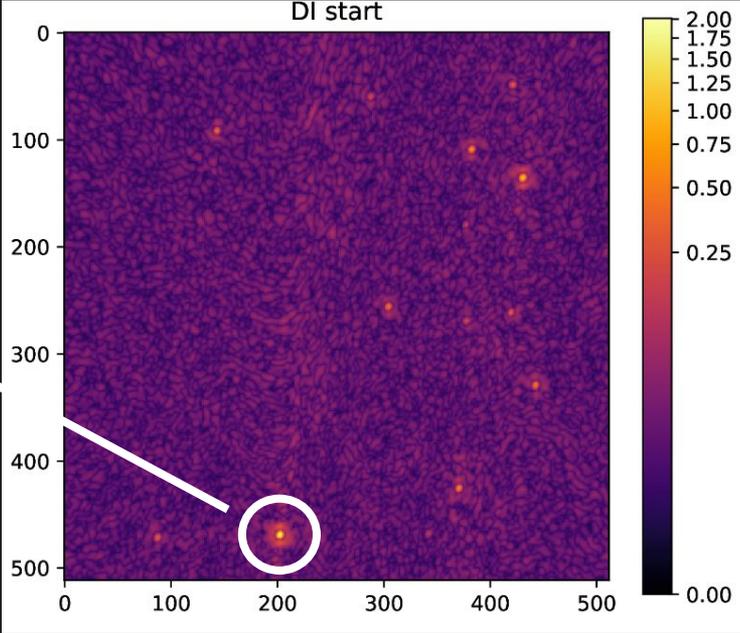


Process sky chunks

Amplitude & Phase

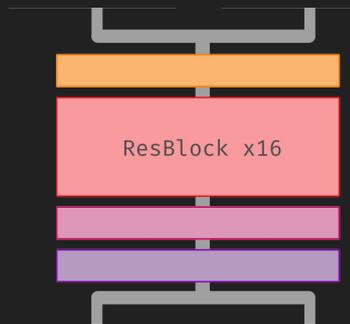


Model Visibilities
(Point sources)



Process sky chunks

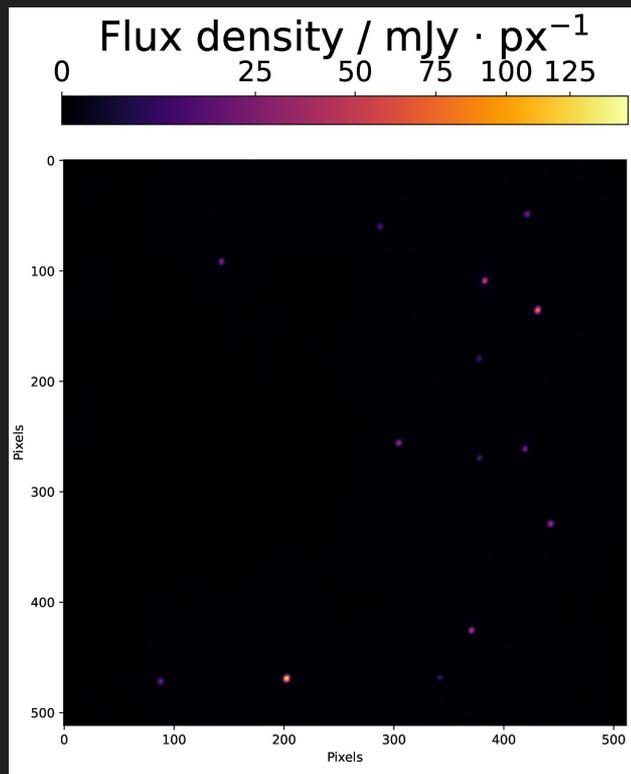
Amplitude & Phase



Model Visibilities
(Point sources)



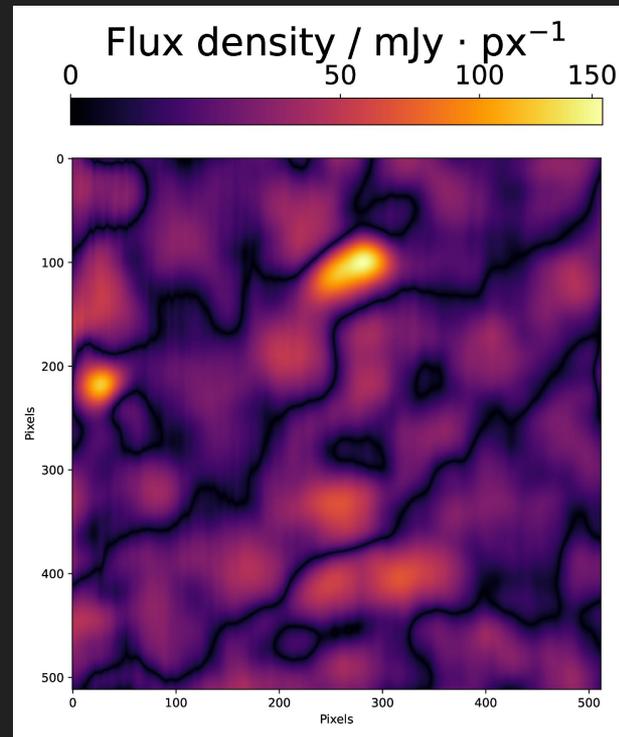
Point source sky model



Process sky chunks

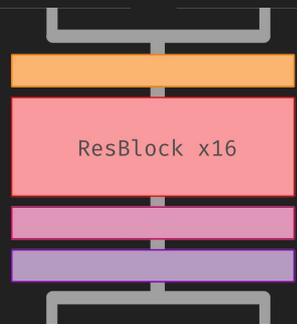
Visibilities — Model Visibilities
(Point sources) →

Diffuse residual



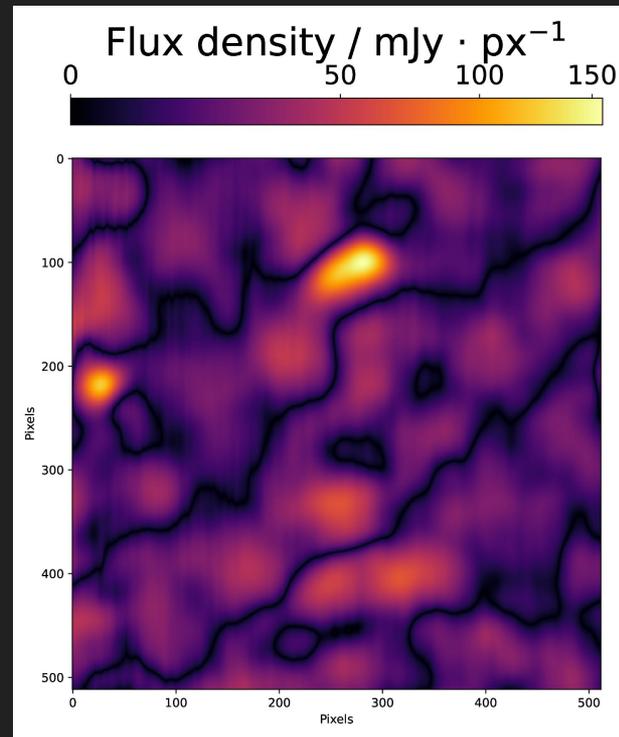
Process sky chunks

Amplitude & Phase



Model Visibilities
(Diffuse)

Diffuse residual



Process sky chunks

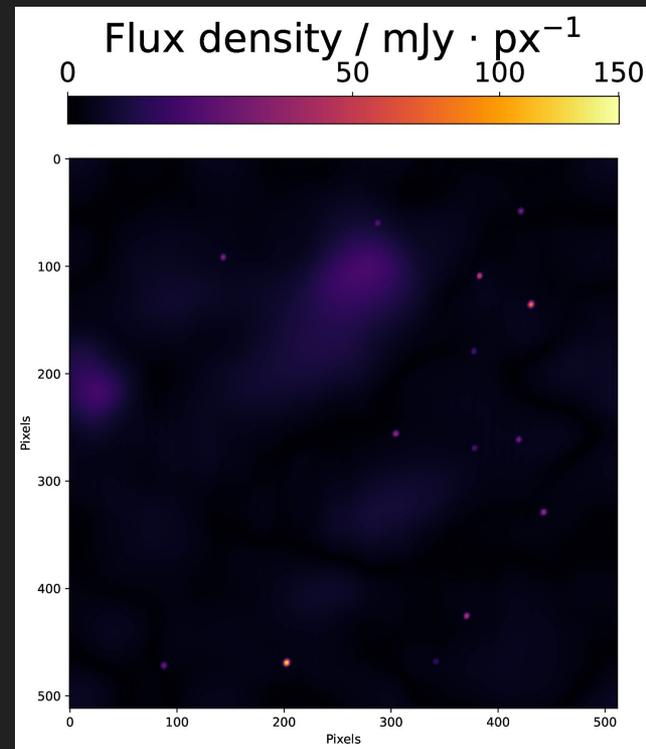
Model Visibilities
(Point sources)

+

Model Visibilities
(Diffuse)



Reconstructed sky



P

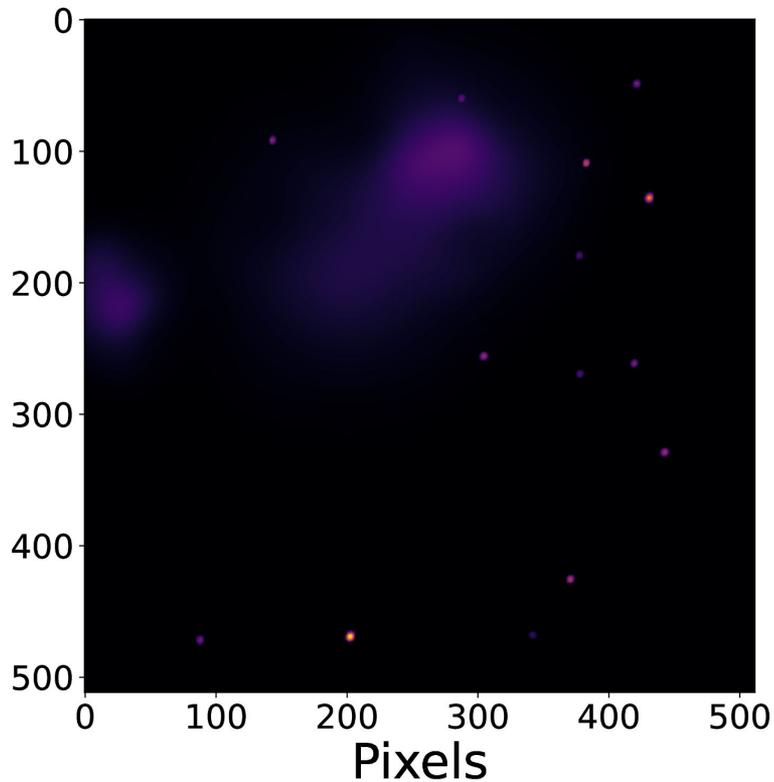
Flux density / mJy · px⁻¹

0 25 50 75 100 125



M

Pixels



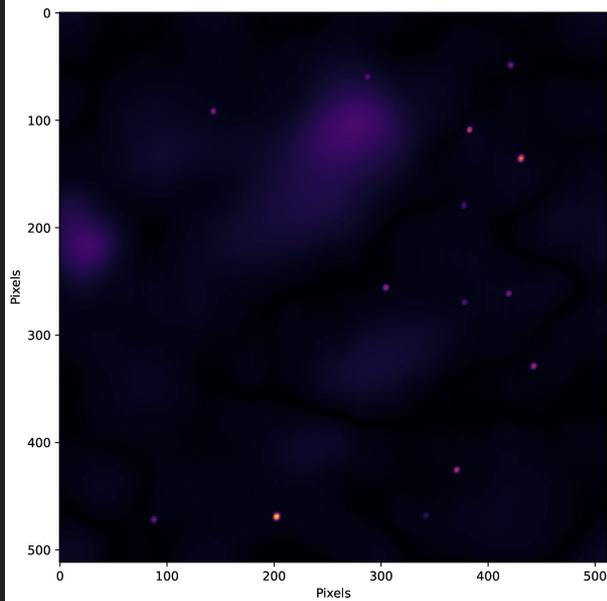
Reconstructed sky

Flux density / mJy · px⁻¹

0 50 100 150



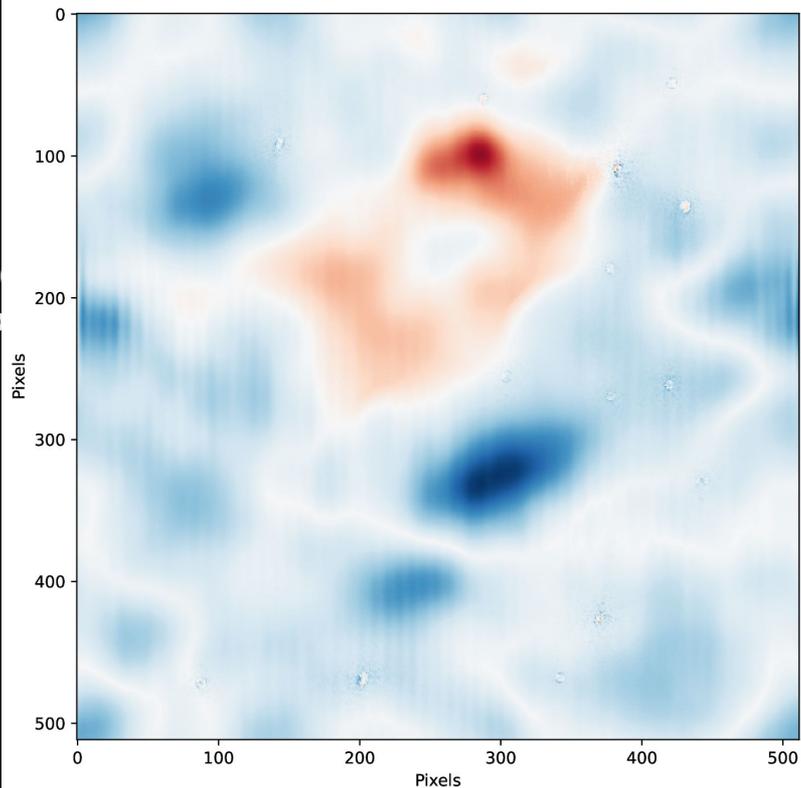
Pixels



P

Flux density / $\text{mJy} \cdot \text{px}^{-1}$

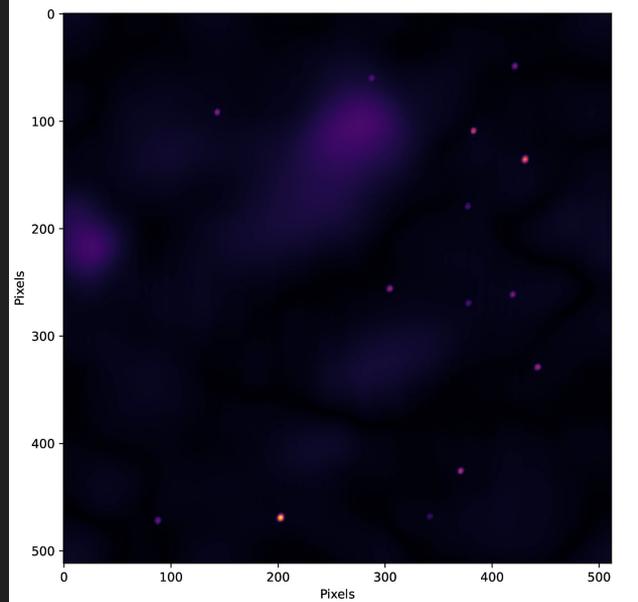
-2 0 2

M
(P

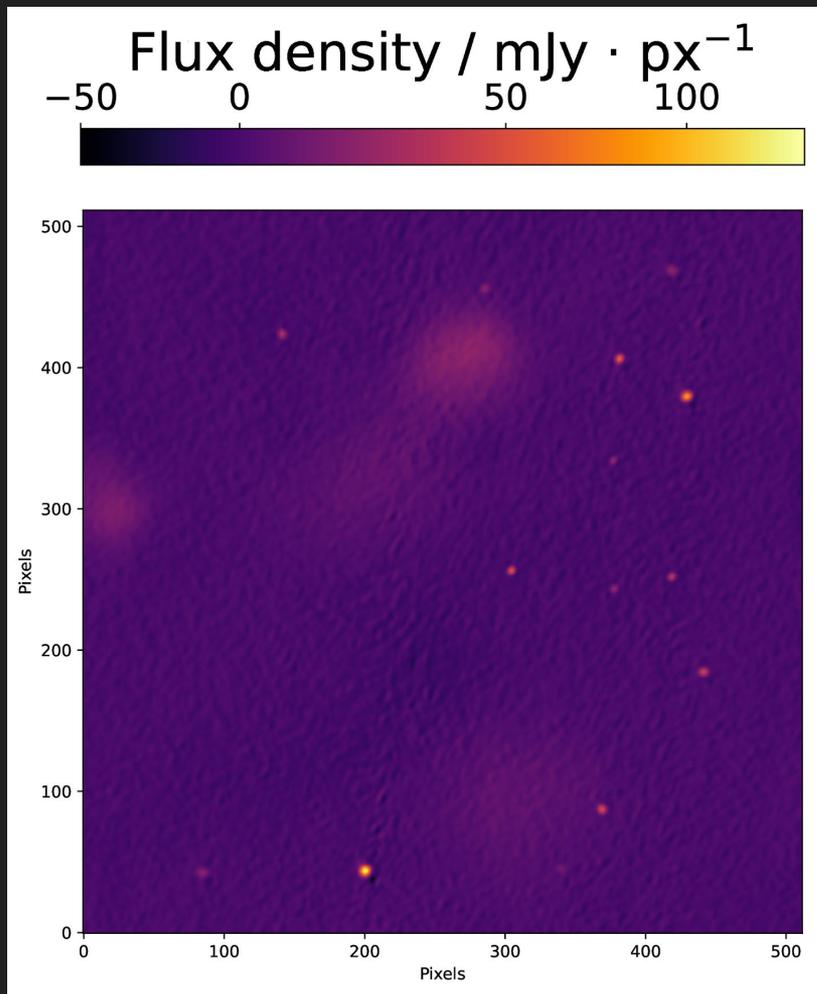
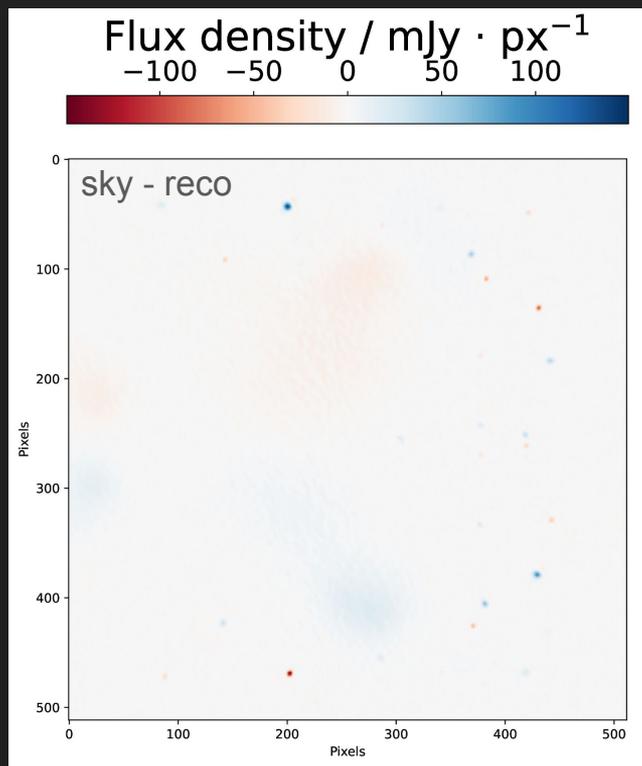
Reconstructed sky

Flux density / $\text{mJy} \cdot \text{px}^{-1}$

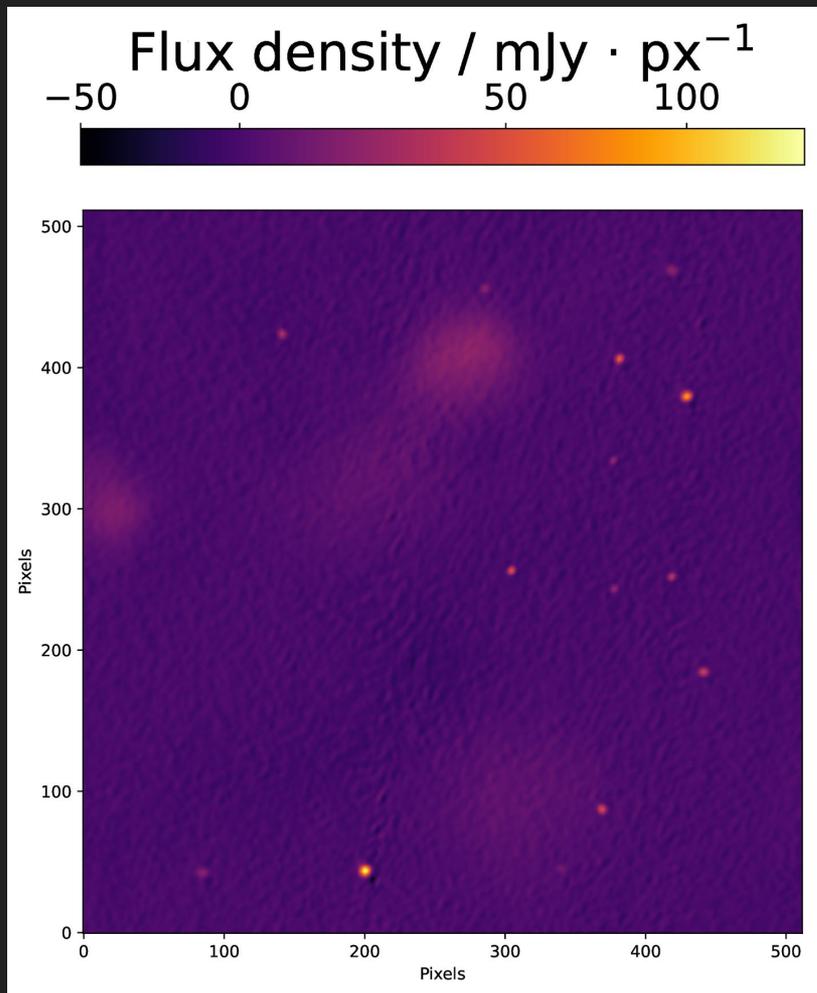
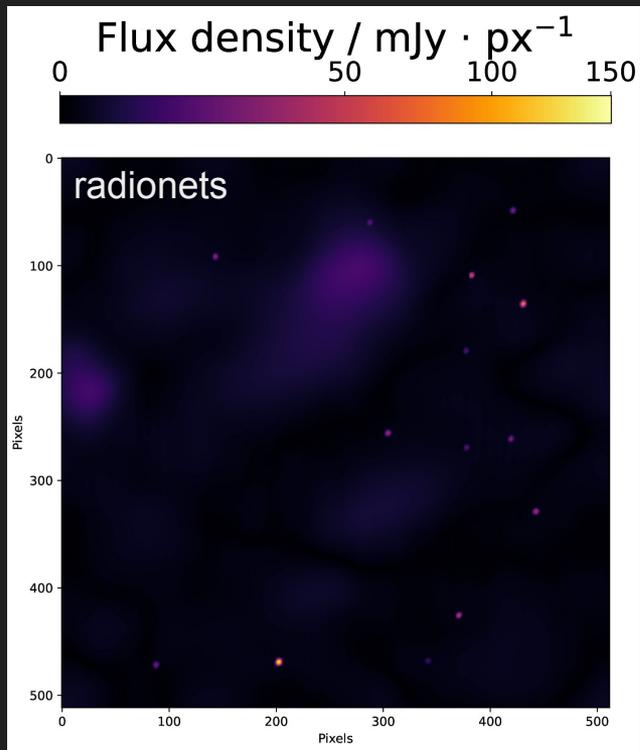
0 50 100 150



WSCLEAN

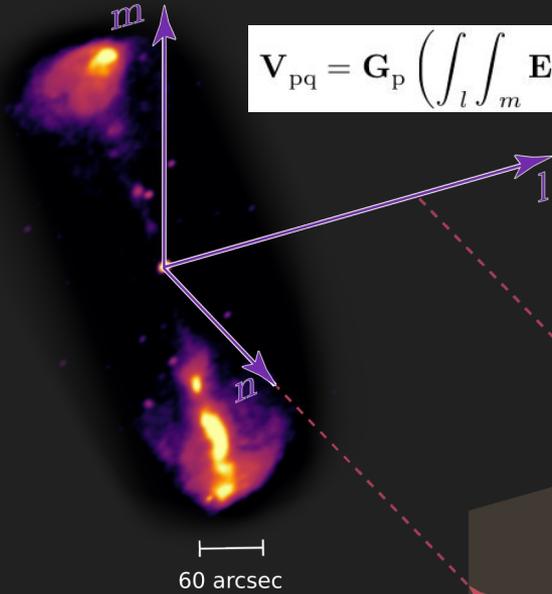


WSCLEAN

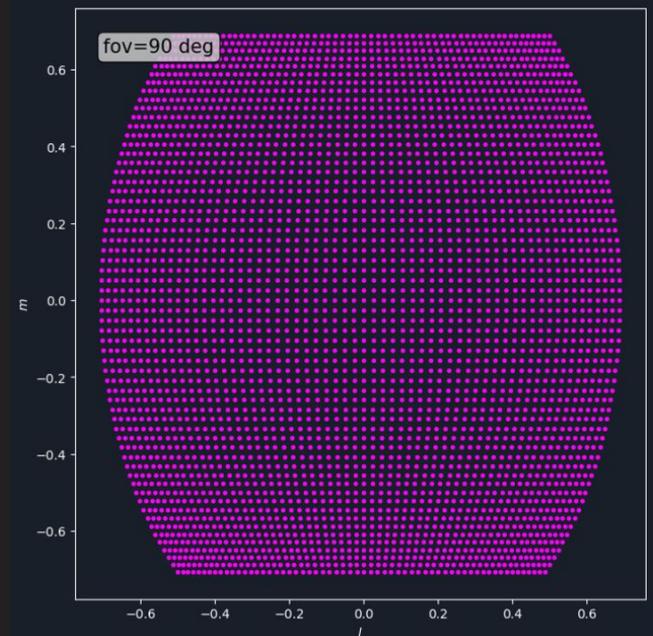
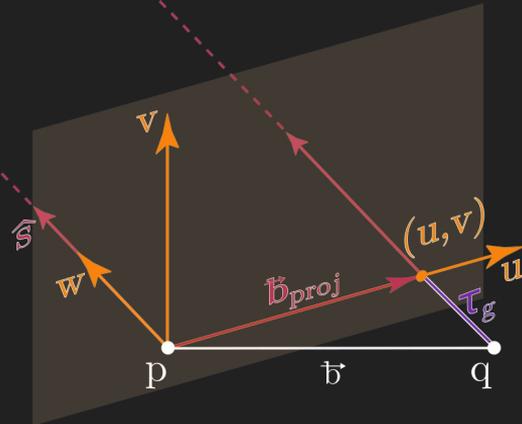


BACKUP

Visibility simulations



$$\mathbf{V}_{pq} = \mathbf{G}_p \left(\int \int_m \mathbf{E}_p(l, m) \mathbf{B}(l, m) \mathbf{E}_q^H(l, m) \exp(-2\pi i(u_{pq}l + v_{pq}m)) dl dm \right) \mathbf{G}_q^H$$



Custom Fourier kernel

Full control

Non-uniform coordinates included

Slow -> Speed-ups

Now:
 ≈ 1 M Vis. in 5 minutes

$$\mathbf{V}_{pq}(l, m) = \sum_{l, m} \mathbf{E}_p(l, m) \mathbf{K}_p(l, m) \mathbf{B}(l, m) \mathbf{K}_q^H(l, m) \mathbf{E}_q^H(l, m)$$

Source Distribution:

$$\mathbf{B}(l, m)$$

Phase Delay:

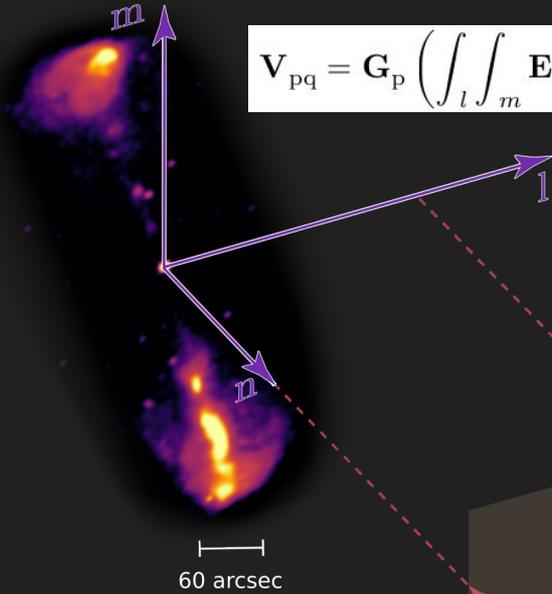
$$\mathbf{K}(l, m) = \exp(-2\pi \cdot i \cdot (ul + vm))$$

Telescope Beam:

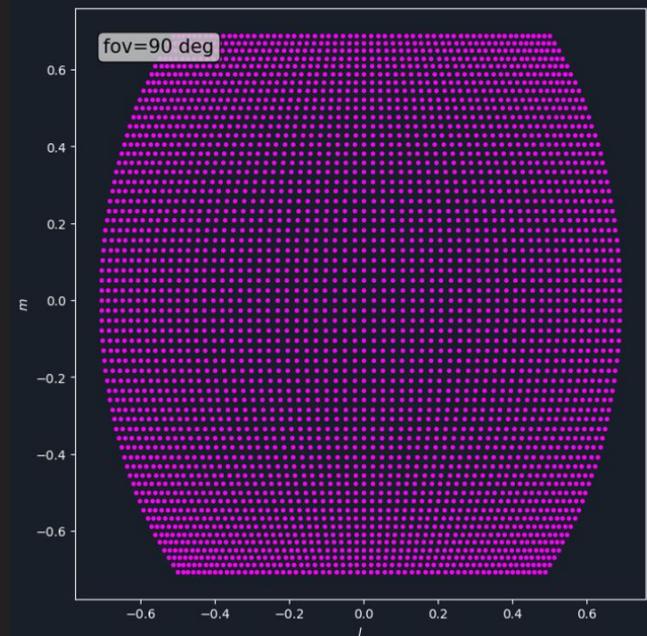
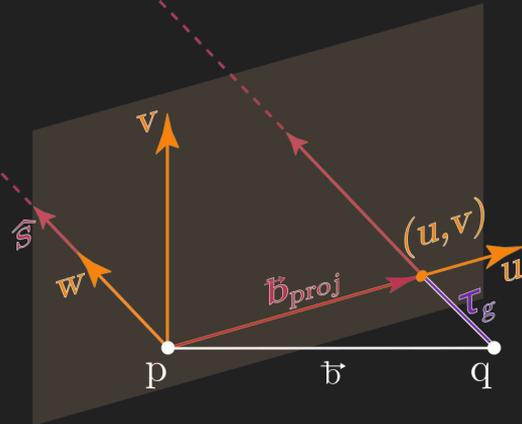
$$\mathbf{E}(l, m) = \text{jinc} \left(\frac{2\pi}{\lambda_{\text{obs}}} d \cdot \theta_{lm} \right)$$

$$\text{jinc}(x) = \frac{J_1(x)}{x}$$

Visibility simulations



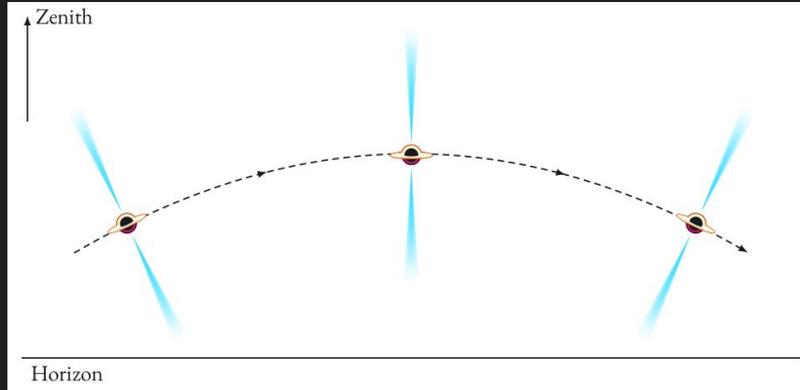
$$\mathbf{V}_{pq} = \mathbf{G}_p \left(\int \int_l \int_m \mathbf{E}_p(l, m) \mathbf{B}(l, m) \mathbf{E}_q^H(l, m) \exp(-2\pi i(u_{pq}l + v_{pq}m)) \, dl dm \right) \mathbf{G}_q^H$$



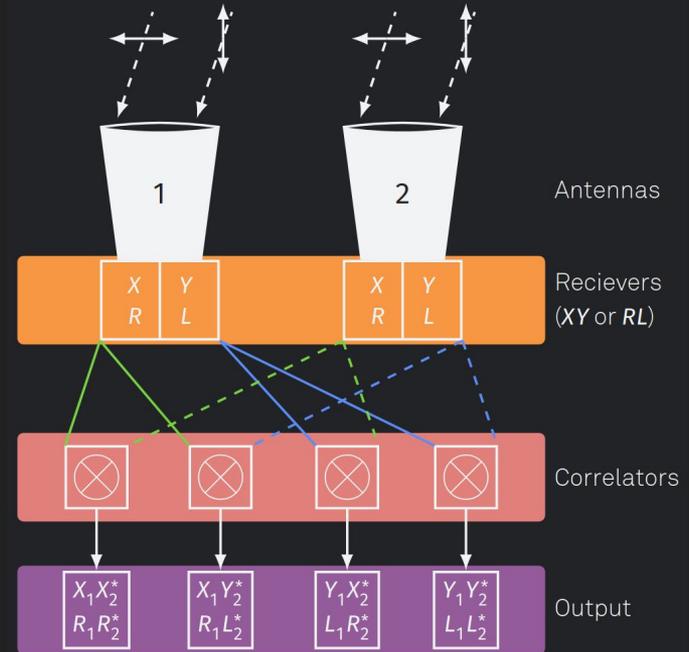
pyvisgen

Polarization:
linear feed
circular feed

Parallactic angle rotation



Simulations



Deep learning model: image inpainting

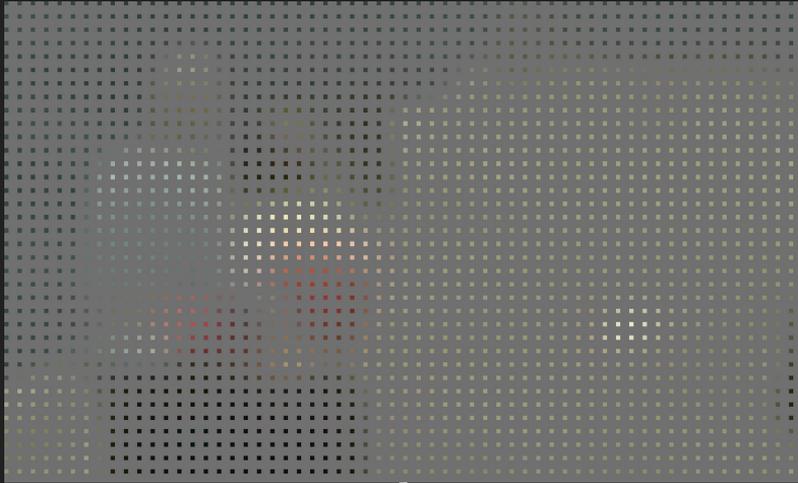


Zoom



(He et al. 2015; Gross & Wilber 2016)

Deep learning model: SRResNet



SRResNet (He et al. 2015; Gross & Wilber 2016)

Deep learning model: Convolutional inpainting

